Final Report

A Complete UV Atlas of Standard Stars

Contract S-57791-Z

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Contents

- 1. Project Summary
- 2. Obervations and Data Retrieval
- 3. Data Reduction with A New IRAF Tool
- 4. Atlas and Spectra of the UV Standard Stars
- 5. Web Site
- 6. Related Publications
- 7. Appendix:
 - o The Atlas of Stars (sorted by spectral type)
 - o Example of Spectra
 - o The Spatial Distribution of the 476 Normal Stars in the Atlas

1. Project Summary

The general objective of this project is to provide a comprehensive ultraviolet spectral atlas of stars based on the data in the IUE Final Archive. The atlas and the project information are presented on our web site.

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http://sol.stsci.edu/~jinger/iue.html
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It links to the Multi-Mission Archive at the Space Telescope Science Institute (MAST). Users can also access this site through

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http://archive.stsci.edu/prep_ds.html.
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The data have been uniformly processed by the IUE NEWSIPS pipeline system. The NEWSIPS gives an improved signal to noise ratio for the data that will be especially beneficial to low signal portions of the spectra; and it also gives more accurate fluxes for stars by adopting the latest absolute calibration and measured instrument parameters. Our new atlas, therefore, provides more uniform and accurate data than the version previously published.

In addition, considerable efforts were devoted to obtain reliable spectral types, V, B-V, and E(B-V). Our goal is to provide the information which can be used with reasonable confidence by scientists.

This comprehensive ultraviolet spectal atlas consists of two sub-atlases. One for "normal" stars, and the other for subluminous stars. The stars in each subgroup are listed in two tables, one sorted by spectral type and the other by HD number or RA (Right Ascension).

- Standard Star Atlas: 476 "normal" stars.
 - o Standard Star Atlas, sorted by Spectral Type
 - o Standard Star Atlas, sorted by HD Number
- Subluminous Star Atlas: 38 subdwarfs and white dwarfs.
 - o Subluminous Stars Atlas, sorted by Spectral Type
 - o Subluminous Stars Atlas, sorted by Right Ascension

The graphics (plots) and flux-wavelength tables of the stars can be accessed via our web page. For each observation, the output resulted from our data reduction includes a pair of spectral graphics, a gif file and a ps (PostScript) file and a w_f (wavelength_flux) text table. A combined spectrum is available, if the observation is a Double Aperture Exprosure or the star has more than two observations.

The applications software, a new MXLO data reduction tool has been developed for our project under the IRAF (Image Reduction and Analysis Facility) environment. It facilitates reducing and analyzing data, generating the graphics and wavelength_flux table for MXLO spectra, retrieving information from various related catologues, and so on. This software is available upon request.

2. Obervations and Data Retrieval

The observations, assembled here, were done mostly by Wu and his collaborators. Only a small number of images were taken from the IUE calibration programs and the other Guest Observers. The observations provide a reasonably comprehensive coverage of the HR diagram. The distribution of the "normal" stars in the atlas is shown in Table I.

The IUE NEWSIPS Low-Dispersion Merged Extracted Image (MXLO) were retrieved from the IUE Final Archive supported by the Multi-Mission Archive at STScI (MAST).

The features of the NEWSIPS MXLO Data are listed below.

• FITS Table

The IUE NEWSIPS MXLO file is a 3-D FITS table, a binary table with fixed-length floating point vectors to contain the extracted fluxes and associated flags.

Data Points

The number of extracted points is always 640. Vacuum ultraviolet wavelength are linearly sampled to a uniform step size.

Aperture

Double aperture low-dispersion spectra will contain two rows, with one row for each aperture.

Absolute Calibration

It converts to absolute fluxes in the range of 1150-1980 Å for short-wavelength spectra and 1850-3350 Å for long-wavelength spectra.

Data Quality Flags

These flages indicate abnormal conditions in the data which can range from fairly minor to quite serious situations. The flags for the MXLO data should be examined carefully in order to ascertain whether or not a particular data point is good or bad. A flag value of zero indicates that there is no know problem associated with the data. In general, flag values of -8 or more negative are indicative of unreliable data.

Small Aperture Spectra

The absolute fluxes for small-aperture data are significantly less reliable than those of large-aperture data. Because centering errors in the small aperture can lead to large variations in the overall observed flux level for individual spectra, it is impossible to determine an absolute small aperture/large aperture ratio (S/L). Therefore, the average of S/L over all wavelengths is normalized to unity. As a result, the small-aperture fluxes are known in a relative sense but not in an absolute one.

Table I. The Distribution of Atlas Stars in Spectral Types and Luminosity Classes

O Stars	VI	v	IV	III	II	Ib	Iab	Ia	If	In_	m
3		2									
4		2									
5		1		3							
6				2			}		1	1	
6.5								1			
7		1		2	2						
7.5				2	1			1			
8		2						2			
8.5		1					1				
9		1	1								
9.5		2		2		1	1	1			
9.7							1				

B Stars	VI	V	IV	III	II	Ib	lab	Ia	If	In	m
0		1		1	1	1					
0.5		1	1	3		1		1			
1		3		2	3	3	2	1			
1.5		2	2	3							
2		2	3	3	1	1		1			
2.5		4	1	2		1					
3		4	2	1	1		1	1			
4		3	2	3		i					
5		3	2	3	1	2		1			
6		3	2	2		1		1			
7		4		1							
7		1		2	2						
8		2		4	1						
9		2	2	1			2	1			
9.5		5		1		-					

A Stars	VI	V	IV	III	II	Ib	Iab	Ia	If	In	m
0		4	1	2	1	1		1			
1		5	1					1			1
2								1	-		
3		6		2	1						1
4		1		-							
5		2	1	1	1	l					
6			1					1			
7		1	4		1		† 				
8		1				1					·
9		-		1			 				

F Stars	VI	V	IV	III	II	Ib	lab	Ia	If	In	m
0		2	1	2		1	l				
1				2	2					<u> </u>	
2		2	1	3		1	 	1			
3		2	1	2				2		·	
4		2	2	2							
5	1	6	2	2		2					
6		7	3	1		1					
7		7	2			2					
8		6	4			2	1	1			
9		7					 				

G Stars	VI	V	IV	III	H	Ib	lab	Ia	If	In	m
0		10	4	2	1	2					
1		5	1_		1	1					
2		8	1			2		1			
3		2			1	1					
4				2	1	1					
5		5	1	2	1	1					
6		2									
7				2							
8		7	3	7	1	1					
9				3	1						

K Stars	VI	V	IV	III	II	Ib	Iab	Ia	If	In	m
0		4	1	11		2					
1		3	3	2	1	1					
2		2	1	9	3	1					
3		2		4	2	1					
4				4		1					
5		1		3			1				
7		1		4		1					

M Stars	VI	V	IV	III	II	Ib	Iab	Ia	If	ln	m
0				4							
0.5		1		1							
1		2									
2				4		1		1			
3		1		4	1						
3.5				1							
4					2						
4.5				2							
5				1							
6				1							
7				1							

3. Data Reduction with A New IRAF Tool

We assumed that the IUE MXLO data have been well-calibrated by the IUE NEWSIPS pipline. And no additional calibration for MXLO data is made by this project.

The run sequence to genaerate the spectral plots and flux tables is :

• Convert Data Format

- O convert the IUE spectra from the MXLO format to the 3D FITS STSDAS table formats
- O convert 3D FITS table to 2D STSDAS table, including the following table columns.
 - * w : wavelength
 - * fL: flux of large aperture image
 - * fS: flux of small aperture image
 - * qL : data quality flag for large aperture image
 - * qS: data quality flag for small aperture image
 - * fLSc: averge combined flux of the large aperture and the scaled small aperture
 - * qLSc: data quality flag for the averge combined image

Make Spectra

O wavelength ranges:

* LWR: 1850 - 3200 Å * LWP: 1850 - 3200 Å * SWP: 1150 - 1980 Å

- O extract ralated information from image header and transfer it to graphics and tables.
- O make flux table
- O make graphics
- O combine images: utilized the STSDAS gcombine as a basic task to make image combine/merge. See *Image Combine* section in our web site for algorithms and parameter specifications. For the atlas, selected high quality spectra of a star were merged to increase the S/N, and to repair spectral regions contaminated by reseaux, cosmic ray hits and other blemishes.
 - * Large Aperture Spectra
 - * Double (Large and Small) Aperture Spectra

4. Atlas and Spectra of the UV Standard Stars

Atlas

This section gives the information for individual stars, including the HD number, star name, spectral type and luminosity class, reference key number for spectral_type and the luminosity_class, RA (Right Ascension) and Dec (Delination), V magnitue, remarks, B-V, and E(B-V).

UV Spectra

This section presents the Spectral Plots and Flux Tables of the UV spectra of the IUE Standard Stars. The information for the star, appeared on the top of the page, is transferred from the Atlas section.

For each observation, the output of the data reduction includes a pair of spectral graphics (a gif file and a ps file) and a w_f (wavelength_flux) text table. Users can access these files by pointing to any of the (gif), (ps) or (w_f) files.

Description of the MXLO Image Log

- O Image: camera(LWR/LWP/SWP) plus image sequence number
 - * (gif) graphics of mxlo spectrum in GIF format
 - * (ps) graphics of mxlo spectrum in PostScript format
 - * (ps) graphics style (color, pattern, and marker)

MXLO Data	Large Aperture Spectra	Small Aperture Spectra
Spectra	Black Solid Line	Violet Dashed Line
Bad Data Points	Red Cross	Blue Asterisk

- O (w_f) mxlo wavelength_flux table in ASCII text format
 - * w : wavelength
 - * fL: flux of large aperture image
 - * qL : DQF for large aperture image
 - * fS: flux of small aperture image
- O Ap : aperture (Large/Small)
- O N: number/T
 - * number the number of exposures along the major axis of the large aperture
 - * T trail mode observation
- O Exp: exposure time in seconds
- O E/C/B: DN (Data Number) level of E/C/B, comments by the IUE staff
 - * E maximum Emission
 - * C maximum Continuum

* B - mean Background

Combined Image

O Single Aperture Exprosure & Double Aperture Exprosure

* Single Aperture Exprosure:

A weighted combined spectrum is given for each camera, if two or more Large Aperture images are available. The weighting factor is the total exposure time.

* Double Aperture Exprosure:

The "pixel-wise" weighting scheme is adopted when combining the Large Aperture(LAp) and the Small Aperture(SAp) data. The weight of a data point is the reciprocal of the variance sigma(i)**2 at that pixel. The sigma is the scaled error taken from the input error map.

- For LAp : error map = sqrt (sL)
- For SAp : error map = sqrt (sS * med)

where sL and sS is the MXLO "sigma" spectrum for LAp and SAp, respectively. The "med" is the median scaled factor of fL/fS over the selected range. And the "sqrt" here is the square root oprator. The image scaling must be done prior to the image combining stage, since only the relative fluxes are available for small aperture.

- O Description of the Combined Image Log
 - * (gif) combined spectrum in GIF format
 - * (ps) combined spectrum in PostScript format
 - * (w_f) combined wavelength_flux table in ASCII text format

Table Column	Single Aperture Spectra	Double Aperture Spectra
wavelength	w	w
flux (of combined spectra)	fcomb	fLSc
DQF (of combined spectra)	qcomb	qLSc

5. Web Site

The contents of the web site of this project is given below. It links to the Multi-Mission Archive at the Space Telescope Science Institute (MAST). Users can also access this site via MAST at

http://archive.stsci.edu/prep_ds.html

Project Information

- About This Project
- Observations
 - o "Nornal" stars
 - o Star Spatial Distribution
 - o Star Spectral Type Distribution
- Data Reduction
 - o MXLO Data
 - o New IRAF Tools
 - o Image Combine
- Final Products: Catalogues and Spectra
 - o Atlas
 - o UV Spectra
- Acknowledgment

IUE NEWSIPS Atlas

- Standard Star Atlas: 476 "normal" stars
 - o Standard Stars Atlas, sorted by Spectral Type
 - o Standard Stars Atlas, sorted by HD Number
- Subluminous Star Atlas: 38 subdwarfs and white dwarfs
 - o Subluminous Stars Atlas, sorted by Spectral Type
 - o Subluminous Stars Atlas, sorted by Right Ascension

IUESIPS Atlas Addendum I & II

- Atlas Addendum I (service provided by NASA/GSFC ADF)
- Atlas Addendum II (service provided by STScI)

Related Publications & IUE Links

- Related Publications
- IUE Archive at STScI
- NASA ADF IUE
- ESA IUE

6. Related Publications

- Wu,C.-C., Mo,J., Schiffer,F.H.,III, & Crenshaw,D.M. 1998, "Ultraviolet Astrophysics Beyond the IUE Final Archive", ESO Conference and Workshop Proceedings SP-413, pp.751-753. "A Comprehensive Ultraviolet Spectral Atlas of Standard Stars"
- Wu,C.-C., Schiffer,F.H., III, & Crenshaw, D.M. 1997, submitted to NASA IUE Newsletter "The IUE Ultraviolet Spectral Atlas, Addendum II"
- Wu,C.-C., Reichert,G.A., Ake,T.A., Crenshaw,D.M., Holm,A.V., Imhoff,C.L., Kondo,Y., Mead,J.M., & Shore,S.N. 1992, NASA Reference Publication 1285. "International Ultraviolet Explorer (IUE) Ultraviolet Spectral Atlas of Selected Astronomical Objects"
- Wu,C.-C., Crenshaw,D.M., Blackwell,J.H., Jr.,Wilson-Diaz,D., Schiffer,F.H.,III, Burstein,D., Fanelli,M.N., & O'Connell,R.W. 1991 NASA IUE Newsletter, No. 43. "IUE Ultraviolet Spectral Atlas, Addendum I"
- Wu, C.-C., Ake, T.A., Boggess, A., Bohlin, R.C., Imhoff, C.L., Holm, A.V., Levay, Z.G., Panek, R.J., Schiffer, F.H., III, & Turnrose, B.E. 1983, NASA IUE Newsletter, No. 22. "The IUE Ultraviolet Spectral Atlas"

8. Appendix:

- The Atlas of Stars (sorted by spectral type)
- Example of Spectra

04	V((f))	HD	164794
09.5	v	HD	93027
B 5	v	HD	188665
A5	Ib	HD	59612
FO	III	HD	7312
F8	v	HD	193901
G2	IV	HD	2151
K0	IIIb	HD	62509
MO.5	III	HD	14651

• Color Print : The Spatial Distribution of the 476 Normal Stars in the Atlas



IUE Ultraviolet Spectral Atlas of Standard Stars

CSC Computer Sciences Corporation

IUE NEWSIPS Atlas: Atlas of the 476 Normal Stars, sorted by Spectral Type

*O*B*A*F*G*K*M*

.05 0	.21 0.5 .02 0.3 .03 0.2	.04 0.3 .01 0.3	.20 0.1	.20 0.5	16 0.3	.17 0.4	.02 0.3	.20 0.1	.19 0.1	.28 0.0	.06 0.3	19 0.4	.09 0.3	.07 0.2	.03 0.3	.09 0.3		ט ט	1000	29 0.5	08 0.1	24 0.0	27 0.0	08 0.2	03 0.2	13 0.1	15 0.0	07 0.1	21 0.0	14 0.4	.15 0.11	24 0.0	22
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8.10 6.46 7.17	7.38 4.04 5.00	6.20 6.24	7.46	8.24	5.79	7.60	5.37	4.88	6.42	77.8	8.26	7.17	5.10	5.63	4.29	4.90															6.15		
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BET CRV G7 IIIa Ba 0.3: 11 12:31:45.3 -23:07:13 2.65 0.90 -0. BET HER G7 IIIa Ba 0.3: 11 16:28:04.1 21:35:50 2.78 A 0.93 -0. TAU CET G8 VP 2 1:41:44.6 -16:12:00 3.50 A 0.73 -0. G8 V 13 2:12:59.0 64:43:32 8.30 0.67 -0. I3 2:12:59.0 64:43:32 8.30 0.67 -0. RHO1 CNC G8 V 2 13 8:49:37.4 28:31:23 5.95 A 0.86 0. 61 UMA G8 V 0.72 -0.	71	6I VIR	99			3:15:47.	18:02:0	٠.	Ą		0.0
BET HER G/V IIIa 11 16:28:04.1 21:35:50 2.78 A 0.93 -0. TAU CET G8 Vp 2 1:41:44.6 -16:12:00 3.50 A 0.73 -0. G8 V 13 2:12:59.0 64:43:32 8.30 0.67 -0. G8 V 13 7:52:02.6 -1:16:47 7.44 0.73 -0. RHO1 CNC G8 V 13 8:49:37.4 28:31:23 5.95 A 0.86 0. 61 UMA G8 V 2 11:38:25.3 34:29:03 5.32 A 0.72 -0.	٠ ا	BET CRV	G7	Ba 0.3		2:31:45.	23:07:1	9.		9.	0.0
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5 G8 V 13 2:12:59.0 64:43:32 8.30 0.67 -0. 6 W 13 7:52:02.6 -1:16:47 7.44 0.73 -0. 2 RHO1 CNC G8 V 13 8:49:37.4 28:31:23 5.95 A 0.73 -0. 1 61 UMA G8 V 2 11:38:25.3 34:29:03 5.32 A 0.72 -0.	2 :	TAU CET	æ (ď		:41:44.	16:12:0	S.	Ą	۲.	0.0
2 RHO1 CNC G8 V 13 8:49:37.4 28:31:23 5.95 A 0.73 -0. 1 61 UMA G8 V 2 11:38:25.3 34:29:03 5.32 A 0.72 -0.	n 4		æ 6	>:		:12:59.	4:43:3	٣.		9.	0.0
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• of UMA GO V $2 11:38:25.3 34:29:03 5.32 A 0.72 -0.$	7 :	KHOI CNC	8 6 5 6	> ;		8:49:37.	8:31:2	٥.	Ą	ω.	۲.
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;	2 11:50:06.2 13 22:11:55.9	13 8:07:26.	2 19:52:51.	1 5:34:09.	15 11:04:28. 11 13:16:11.	Ba 0.4 11 15:00:03.	LI 18:41:10: 12 20:15:16.	1 22:47:35.	1 21:10:48.	2 6:40:51.	11:04:21.	1 19:12:33.	II 8:52:45.	K Stars	1:44:06.	3 15:07:28.	2 19:32:27. 3 20:12:10.	2 20:44:16.	IV 1:00:30.	2H-1 H&K-0.5 12 0:41:04.	N-2 11 5.49.10	3a 0.1 11 6:45:15.	1 7:42:15.	IN 1 10:08:08.	1 20:44:11.	1 22:47:53.	CN-1.5 11 23:14:34.	1 21:43:46.	7.32.48.	2 4:12:58.	14:53:45.	1 15:49:20.8	J 13:34:30.
5	89 89 f	PSI CNC G8 IV 31 AOI, G8 TV H	. G8 IV		GAM HYA G8- IIIa	- IIIa:	TIII 85 A	III 85	ZET CYG G8 II CN	9 6	11	DRA G9 II	I-II		KO V	KO V	4		KO III-	III	LEP KO III	40N KO III	GEM KO IIIb	LAM HYA KO III C	CYG KO-	CEP KO- III	M PSC K0- III:	7 KO	·	2 ERI K1-	K1	KAP CRB K1 IVA PHI SER K1 TV	

999 0.0 011 -0.0 119 0.1 335 0.2 57 0.3 888 -0.0 02 -0.0	256 0.1 225 0.0 21 -0.0 22 0.0 117 0.0	118 0.0 555 0.4 550 0.3 550 0.2 55 0.0 99	.36 0.10 .23 -0.03 .24 0.014 .54 0.13 .54 0.13 .54 0.05 .47 0.04 .47 0.04 .45 0.05 .67 0.01 .58 0.07 .58 0.07 .58 0.03 .58 0.07
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:31:1 :37:0 :42:0 :15:2 :15:2 :57:1 :37:3 :04:1	0:19:4 2:28:2 8:11:2 3:31:0 6:14:3 6:14:3	52: 52: 10: 38: 46: 53: 38:	5:14:07 34:26:07 16:32:06 -10:56:39 33:05:19 -55:29:06 75:55:05 74:21:35 -43:13:47 77:30:59 38:29:38 38:29:38
	: 59:40. : 29:40. : 36:08. : 49:55. : 41:00.	7:52:39. 3:17:18. 6:43:21. 7:24:01. 1:41:43. 6:04:41.	1:38:49.6 5:24:19.8 14:17:23.1 4:53:43.9 10:15:24.5 17:21:08.2 8:13:48.2 14:27:36.1 14:27:36.1 14:50:49.6 22:03:09.4 9:06:09.3 21:04:40.0 4:33:02.8 15:32:51.2 17:55:26.5 6:21:02.8 21:04:38.0
	11 13 13 1 Ba -1 11 1	1 13 13 13 2	
IV III-IV III CN+2 5 Ib V V V IV III	III III III III CN 1 Ca III	iii II-III IID-IIIa ID V	- IIIb Ba 0.1 III CN+2 III - III CN 2 II IIa IIb-IIa III Ba 0.3 III III III III III III III III III I
+ 1 •	2	2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	XXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXXX
TAU CRB ETA PAV THE HER ZET CEP EPS ERI 62 AUR	TAU GEM SIG HYA MU LEO IOT DRA BET OPH	ALF TRA SIG OPH EPS PEG	NU PSC PHI AUR 20 BOO 18 LIB IOT AUR BET CNC 5 UMI BET UMI NU PEG LAM VEL 61 CYG A ALP TAU THE UMI GAM DRA PSI1 AUR
145148 145148 160635 163770 210745 166620 166620 51440	54719 66141 72184 73471 85503 137759 161096	163588 20644 150798 157999 206778 144872	10380 35620 125560 132345 31398 89388 157244 69267 127700 131873 201091 29139 164058 44537 201092

1.49 -0.04 A 1.73 0.11		A 1.58 0.01	0	0	0	0	0	1.51 -0	0	0	0	0	0	0	~	0	0	0	0	2.07 0	2.33 0	0	1.58 -0	0	1.57 -0	0	0	0	0	0	9	0-	0	0	0
4.83		2.05	3.41	3.05	3.85	8.49	2.73	4.03	4.77	4.10	4.94	5.05	0.48	. 50	8.65	4.80	4.69	4.66	4.83	4.35	4.10	8.90	4.67	2.87	3.38	3.30	4.29	1.62	3.39	4.28	5.26	5.58	4.59	5.00	69.9
11:35:38 -27:51:43		35:21:21	-43:34	41:45	69:36	-0:29	-3:34	6:48	17:40	18:17	21:56	-8:56	7:23	36:18	22:3	19:55	8:17	54:56	48:44	18:33	58:33	59:33	24:51	22:32	3:40	-25:05	45:56	-56:50	38:38	36:50	44:05	23:37	66:07	41:59	-2:26
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K7 III K7 Ib		MO IIIa	H	H	MO III Ca-1	_	MO.5 III	M1 IIIab	M1- IIIb	M1- IIIab	W1 IIIb	il+ Ib-IIa	M1-M2 Ia-Ib	42 V	dM2	42+ III	42- IIIab	12 IIIab	TII Z	12 Iab-Ib	12 Ia	43 V	43 III	43 IIIab	iii -	f3- III	i II	43.5 III	14 IIb-IIIa	14 II	44.5 III	4.5 IIIa	45 III	11 - 111 - 9	17 III:
29 HER ISIG CMA		BET AND						NU VIR			106 HER		ALP ORI			CHI PEG	PI LEO	83 UMA	8 AND	119 TAU	MU CEP		PSI PEG	MU GEM											
149161 52877		6860	9053	89758	100029	111631	146051	102212	112769	141477	168720	49331	39801	95735	216399	1013	86663	119228	219734	36389	206936	173739	224427	44478	112300	133216	40239	108903	19058	175588	123657	145713	132813	148783	207076

--- End of Atlas (Sort by Spectral Type) ---



IUE Ultraviolet Spectral Atlas of Standard Stars

CSC

Computer Sciences Corporation

UV Atlas Home

Project Info

NEWSIPS Atlas

IUESIPS Atlas

IUE Links

4 -> >

IUE NEWSIPS Atlas: Subluminous Star Atlas, sorted by Spectral Type

Star Name	Sp_Type	RA(1950.0) hh:mm:ss	DEC(1950.0) dd:mm:ss	V mag	B-V mag	
HD 49798	sd0	06:46:34.85	-44:15:33.5	8.27	-0.30	
BD +75 325	sd0	08:04:43.44	75:06:47.7	9.54	-0.37	
AGK +81-266	sd0	09:13:42.53	81:56:11.29	11.85	-0.34	
HZ 44	sd0	13:21:19.10	36:23:38.00	11.71	-0.27	
BD +28 4211	sd0	21:48:57.30	28:37:44	10.52	-0.34	
GD 108	sdB	10:00:47.33	-07:33:31.2	13.56	-0.23	
FEIGE 65	sdB	12:33.24	42:39	12.01	-0.24	
FEIGE 66	sdB	12:34:54.69	25:20:30.4	10.51	-0.26	
G 191-B2 B	DA1	05:01:31.3	52:45:50	11.78	-0.29	
GD 71	DA1	05:49:34	15:52.7	13.04	-0.24	
BPM 16274	DA2	00:50:03.18	-52:08:17.4	14.20	-0.02	
GD 50	DA2	03:48:50.06	-00:58:30.4	14.05	-0.28	
G 87-7	DA2	06:44:15	37:34.9	12.04	-0.07	
GD 394	DA2	21:11:03.1	49:53:53	13.10	-0.23	
HZ 2	DA3	04:09:57.00	11:44:13.92	13.86	-0.05	
GRW +70 5824	DA3	13:38:51.77	70:17:08.5	12.79	-0.09	
G 93-48	DA3	21:52:25.33	02:23:24.3	12.74	-0.01	
FEIGE 108	DAS	23:13:36	-02:07	12.90	-0.28	
40 ERI B	DA4	04:13:03.66	-07:44:08.9	9.50	0.11	
HZ 4	DA4	03:52:37.90	09:38:34.08	14.47	0.08	
LB 227	DA4	04:06:36.89	17:00:03.96	15.35	0.05	
G 226-29	DAV	16:47:38	59:08:42	12.19	0.04	
LDS 532-81	DA6	08:39:35.9	-32:46:55	12.0	0.16	
ROSS 627	DA7	11:21:37.9	21:38:05	14.17	0.37	
FEIGE 7	DABP3	00:41:15	-10:16.6	14.52	0.01	
GD 323	DAB	13:02:27	59:42:54	14.52	-0.13	
GD 303	DB4	10:11:17	57:03.5	14.62	-0.14	
L 1573-31	DB4	19:40.24	37:24	14.57	-0.10	
LDS 749 B	DB4	21:29:36.60	00:00:00.00	14.73	-0.04	
LDS 678 B	DBQ	19:17:52.9	-07:45:34	12.33	0.05	
G 175-34 B	DC7	04:26:50	58:53.3	12.44	0.31	
HZ 21	DO1	12:11:24	33:12	14.63	-0.28	
PG 1159-03	DO2	11:59:12.3	-03:28:57	14.87	-0.37	
FEIGE 34	DO	10:36:40.00	43:21:52.00	11.12	-0.30	
FEIGE 110	DOp	23:19:58.39	-05:09:55.8	11.50	-0.30	
VAN MAANEN 2	DZ8	00:46:28.8	05:10:21	12.46	0.52	
L 145-141	DQ6	11:42:58.0	-64:33:34	11.44	0.16	
GRW +70 8247	DXP	19:00:40	70:35.6	13.14	0.08	



jinger@stsci.edu last updated: Feb. 29, 2000

IUE Standard Stars Atlas (NEWSIPS)

HD 164794 9 SGR

 Sp_Type
 RA
 DEC
 V
 Rem
 B-V
 E(B-V)

 04
 V((f))
 18:00:48.4
 -24:21:48
 5.97
 0.03
 0.35

* Low-Dispersion Merged Extracted Image (MXLO)

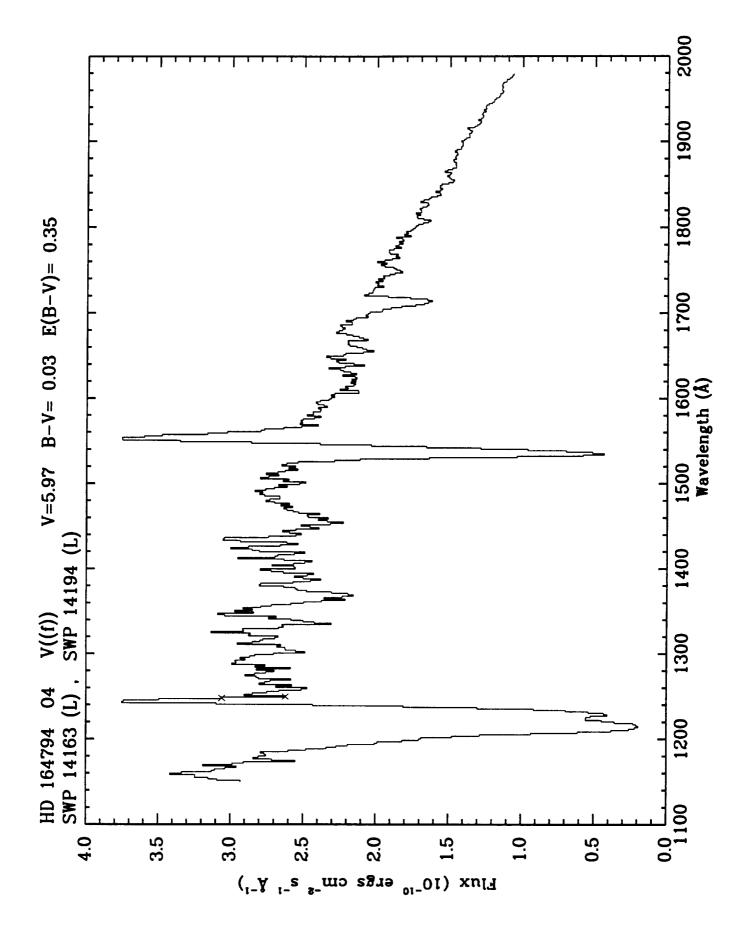
Image	Ap	N	Екр	E/C/B
SWP14163(gif)(ps)(w_f)	L	т	18.280	/230/23
SWP14194(gif)(ps)(w_f)	L	${f T}$	16.121	/226/25
LWR10768(gif)(ps)(w_f)	L	1	3.151	/205/26
LWR10787(gif)(ps)(w_f)	L	T	13.541	/203/25

* Combined Image

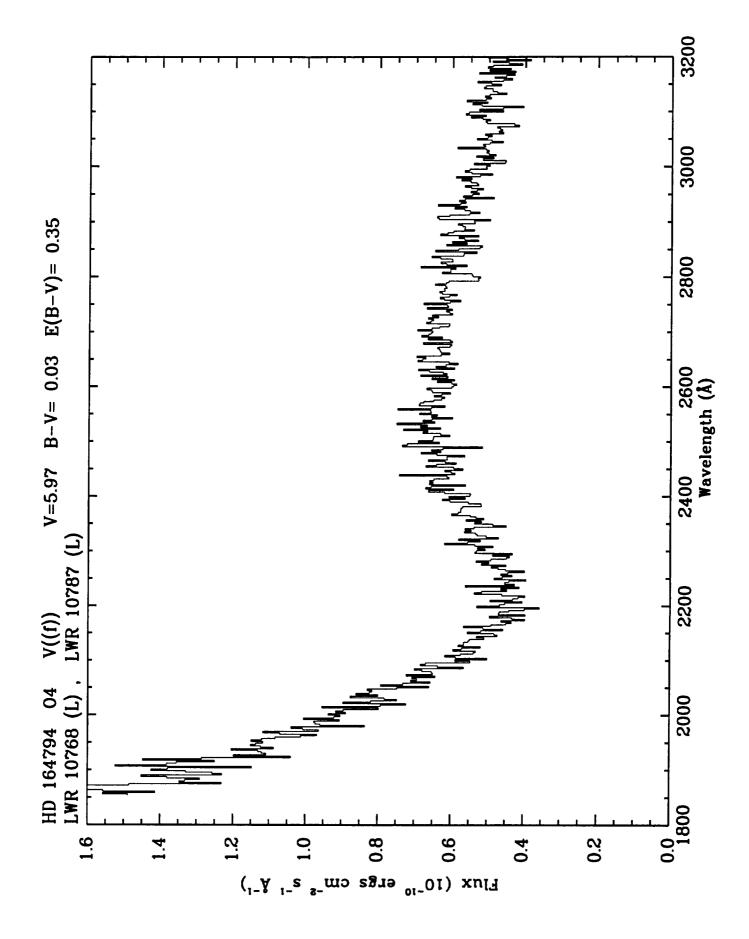
SWP14163,SWP14194 (gif) (ps) (w_f) LWR10768,LWR10787 (gif) (ps) (w_f)

UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	_

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IUE Standard Stars Atlas (NEWSIPS)

HD 93027

Sp_Type	RA	DEC	v	Rem	B-V	E(B-V)
09.5 V	10:41:18.3	-59:52:40	8.72		0.02	0.28

* Low-Dispersion Merged Extracted Image (MXLO)

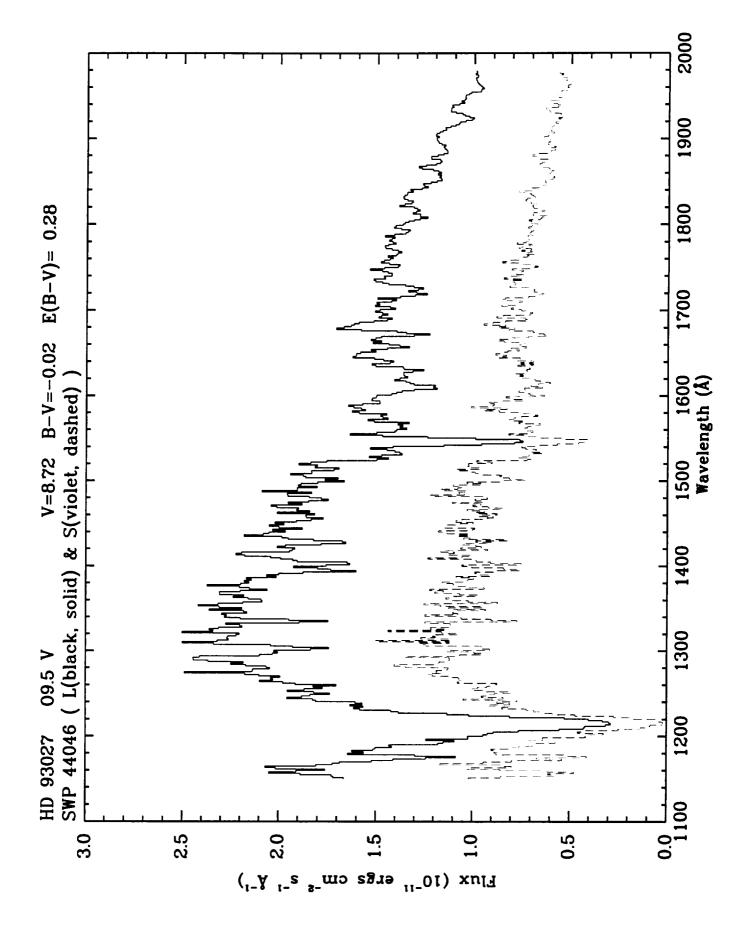
Image	Ap	N	Ежр	E/C/B
SWP44046(gif)(ps)(w_f)	L	т	161.469	/170/18
	S	1	59.672	/140/18
SWP46597(gif)(ps)(w_f)	L	T	216.881	/208/15
LWP22441(gif)(ps)(w_f)	L	\mathbf{T}	131.416	/170/38
	S	1	239.493	/5X/38
LWP24604(gif)(ps)(w_f)	L	T	197.124	/242/35

* Combined Image

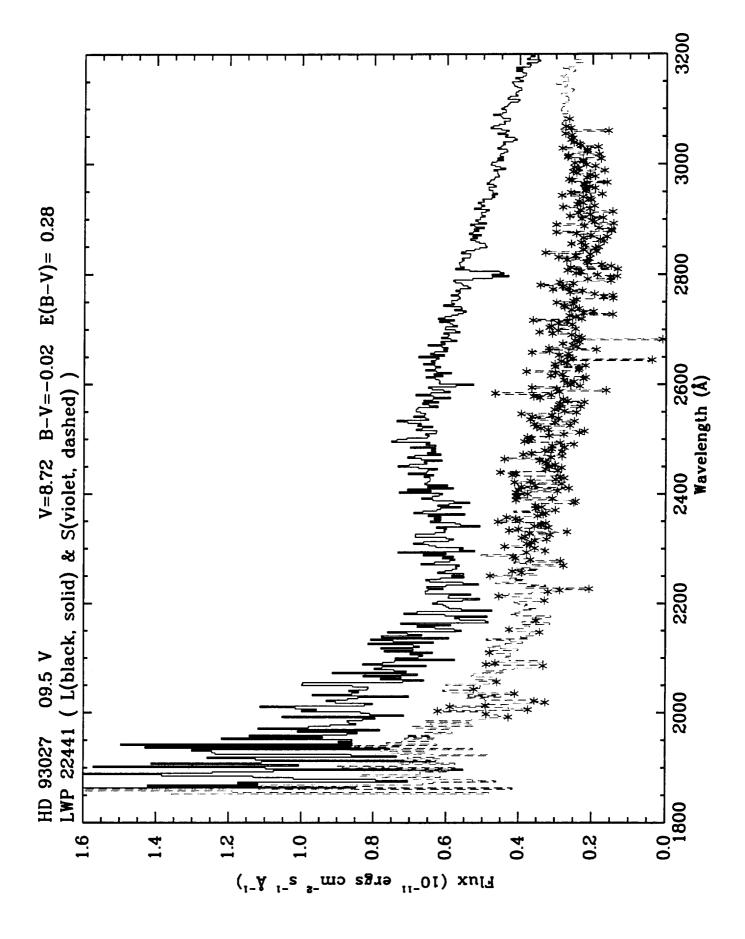
SWP44046, SWP46597 (gif) (ps) (w_f) LWP22441, LWP24604 (gif) (ps) (w_f)

UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	•

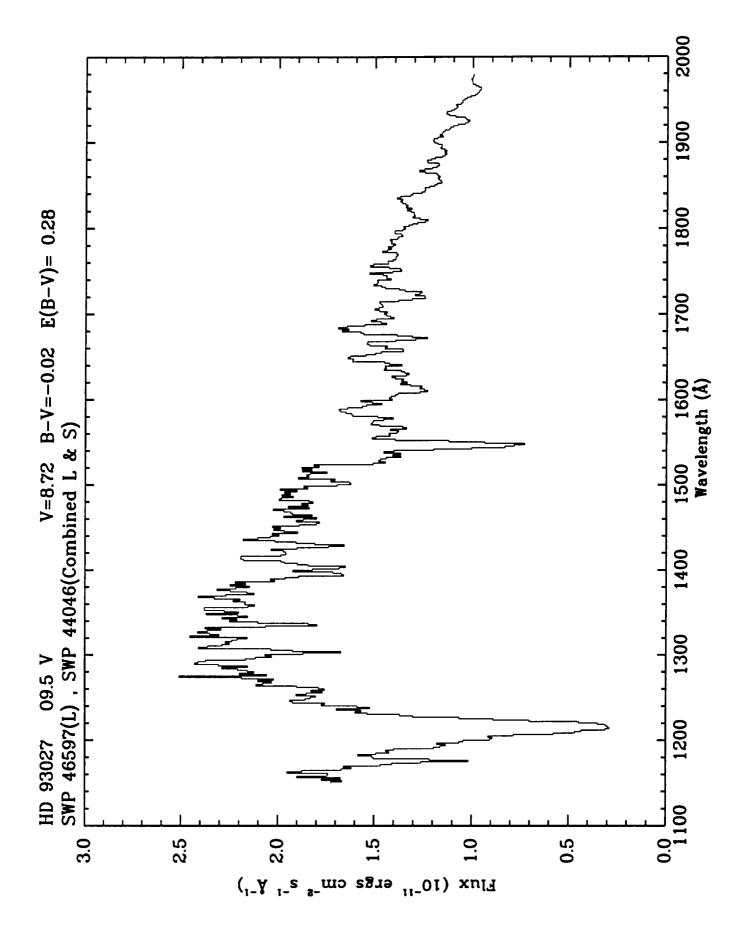
			•

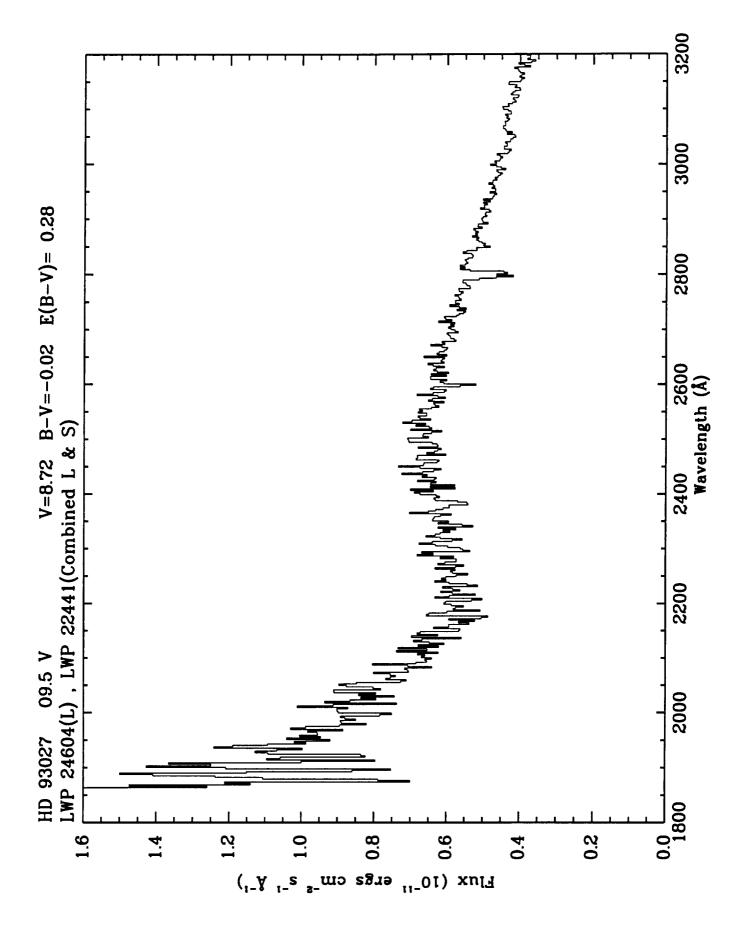


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HD 188665 23 CYG

Sp_Type	RA	DEC	v	Rem	B-V	E(B-V)
B5 V	19:52:15.8	57:23:30	5.14		-0.14	0.02

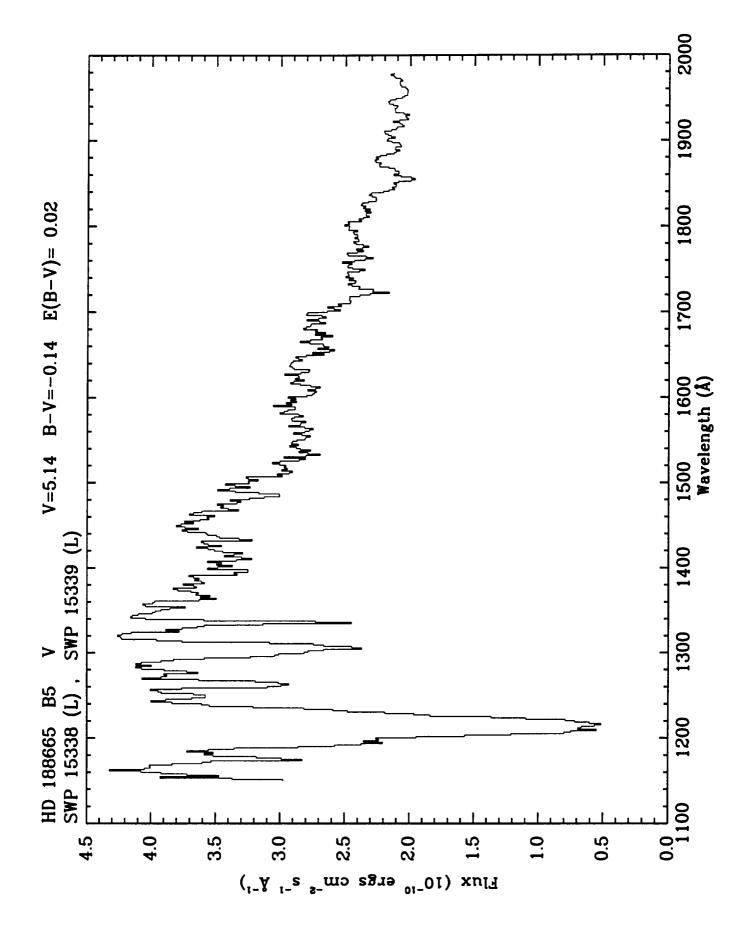
* Low-Dispersion Merged Extracted Image (MXLO)

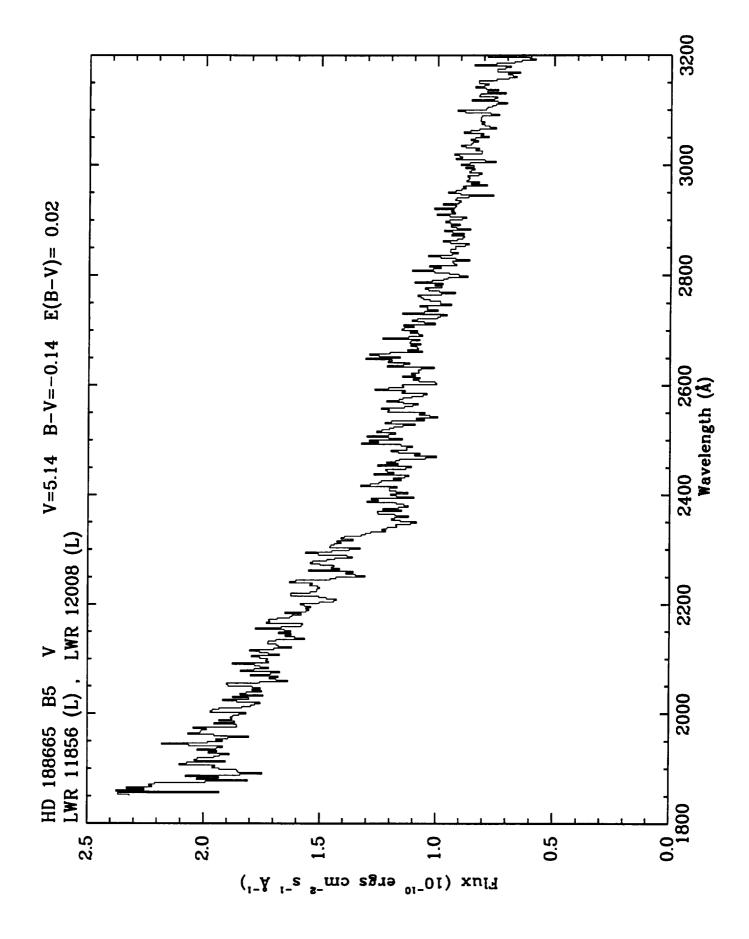
Image	Ap	N	Ехр	E/C/B
SWP15338(gif)(ps)(w_f)	L	T	5.371	/120/25
	S	1	0.705	//
SWP15339(gif)(ps)(w_f)	L	Т	8.055	/180/29
	S	1	1.058	//
$LWR11856(gif)(ps)(w_f)$	L	T	9.023	/208/30
LWR12008(gif)(ps)(w_f)	L	T	10.152	/205/25

* Combined Image

SWP15338, SWP15339 (gif) (ps) (w_f) LWR11856, LWR12008 (gif) (ps) (w_f)

UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	





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HD 59612

Sp_Type RA DEC V Rem B-V E(B-V) A5 Ib 7:27:43.9 -22:55:09 4.85 AB 0.23 0.13								
a5 Tb 7:27:43.9 -22:55:09 4.85 AB 0.23 0.13	Sp_	Гуре	RA	DEC	v	Rem	B-V	E(B-V)
	A 5	Ib	7:27:43.9	-22:55:09	4.85	AB	0.23	0.13

* Low-Dispersion Merged Extracted Image (MXLO)

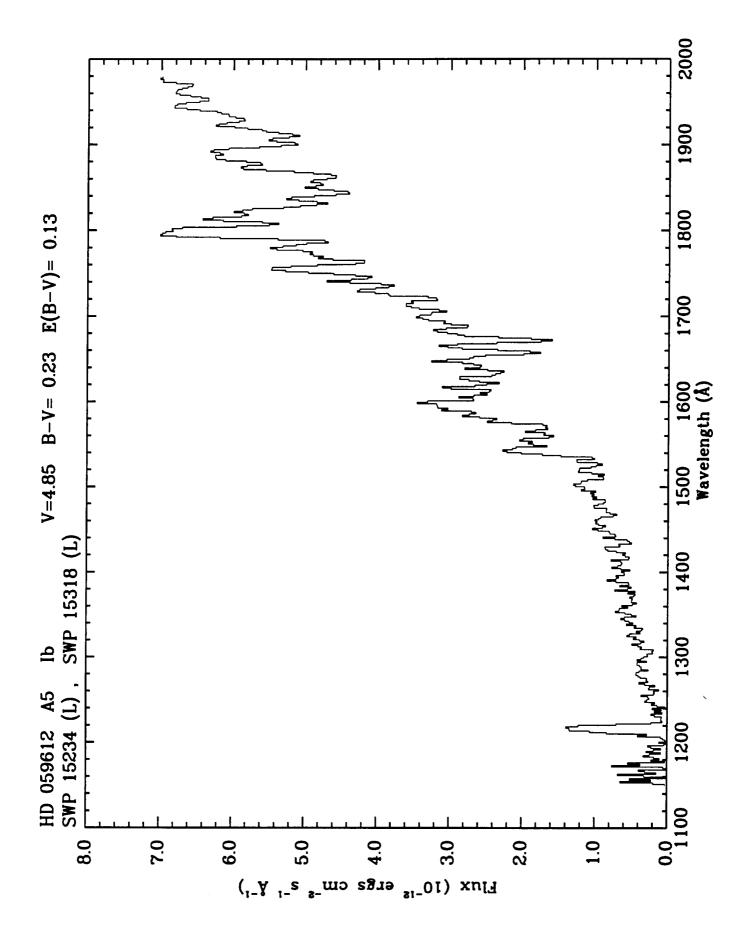
Image	Ap	N	Екр	E/C/B
SWP15234(gif)(ps)(w_f)	L	т	139.848	/100/26
SWP15318(gif)(ps)(w_f)	L	${f T}$	310.199	/180/30
LWR11748(gif)(ps)(w_f)	L	Т	79.135	/180/26
LWR11824(gif)(ps)(w_f)	L	${f T}$	237.749	/2X/28

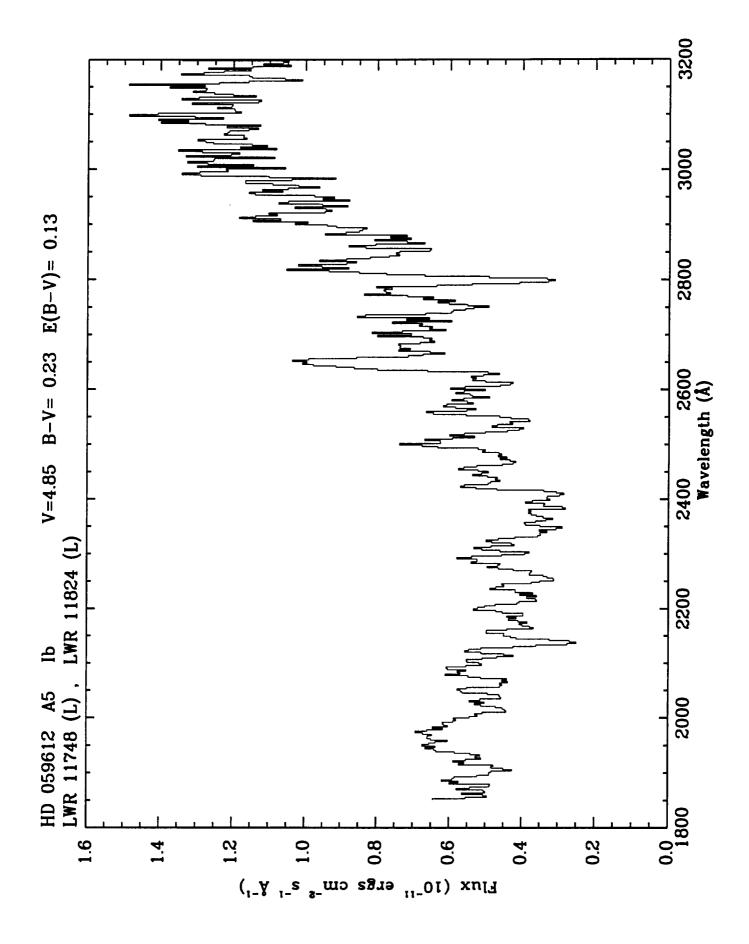
* Combined Image

SWP15234, SWP15318 (gif) (ps) (w_f) LWR11748, LWR11824 (gif) (ps) (w_f)

UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	_	







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HD 7312

Sp_Type	R A	DEC	v	Rem	B-V	E(B-V)
FO III	1:10:27.3	-38:07:15	5.91		0.29	-0.03

* Low-Dispersion Merged Extracted Image (MXLO)

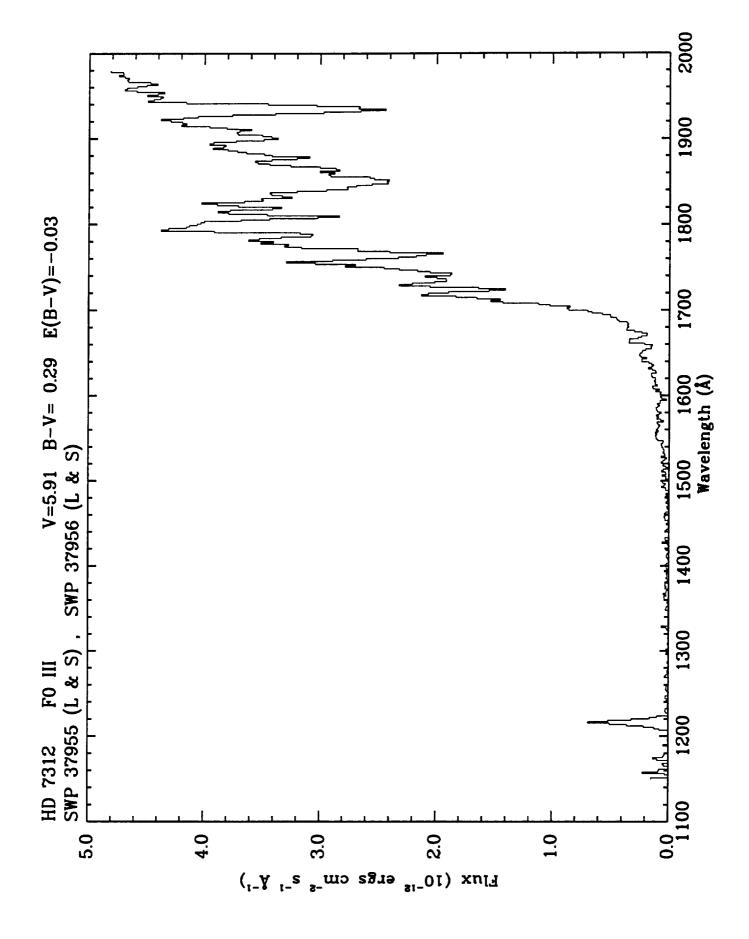
Image	Ap	N	Екр	E/C/B
SWP37955(gif)(ps)(w_f)	L	3	3598.765	/60X/34
	S	1	899.761	/5X/36
SWP37956(gif)(ps)(w_f)	L	3	374.394	/203/21
	S	1	119.473	/107/24
LWP17090(gif)(ps)(w_f)	L	3	418.652	/4X/28
	S	1	119.480	/2X/25
LWP17091(gif)(ps)(w_f)	L	T	65.708	/157/37

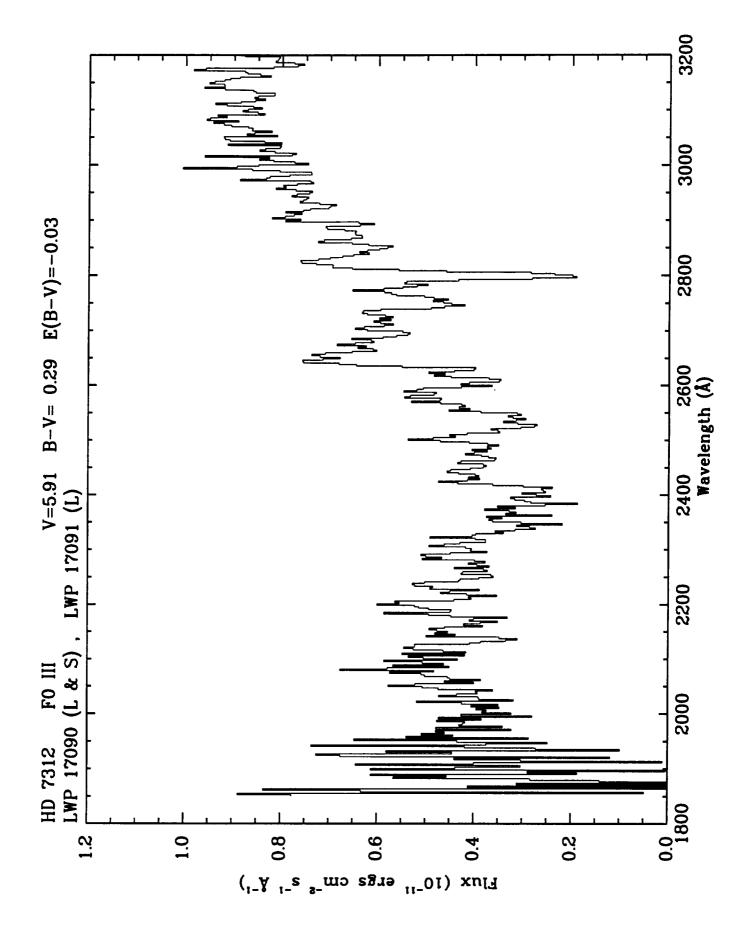
* Combined Image

SWP37955,SWP37956 (gif)(ps)(w_f) LWP17090,LWP17091 (gif)(ps)(w_f)

UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	•

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HD 193901

Sp_Type RA DEC V Rem B-V E(B-V)
F8 V 20:20:38.8 -21:31:05 8.65 0.55 0.02

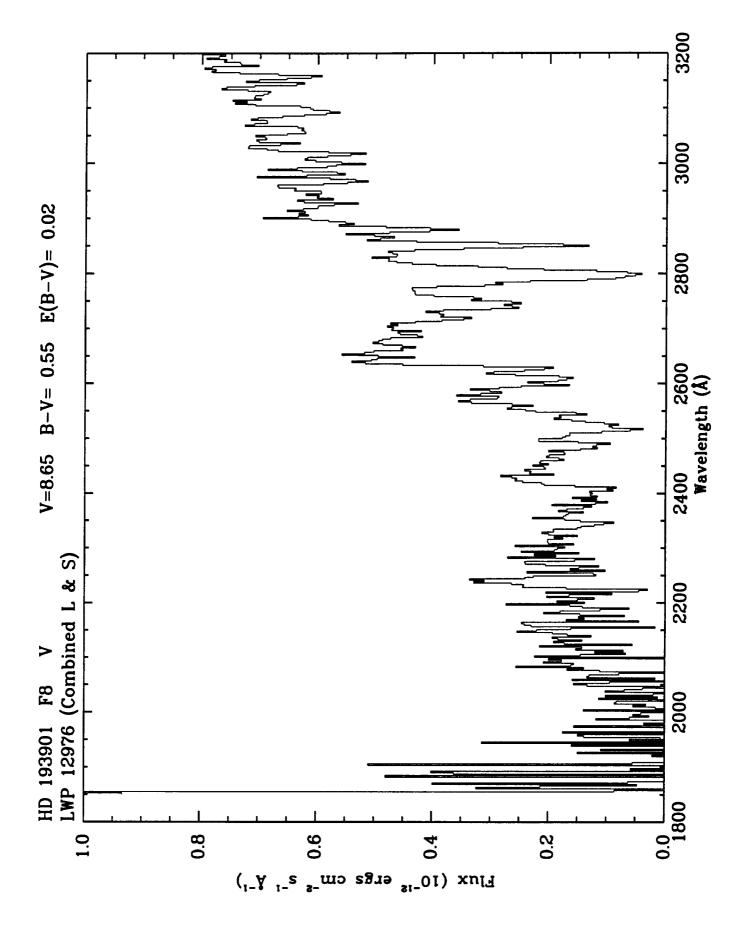
* Low-Dispersion Merged Extracted Image (MXLO)

Image	Ap	N	Ежр	E/C/B
LWP12976(gif)(ps)(w_f)	L S	3		/231/46 //42

* Combined Image

LWP12976(gif)(ps)(w_f)

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UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links		l





HD 2151 BET HYI

Sp_7	(Abe	RA	DEC	v	Rem	B-V	E(B-V)
G2	IV	0:23:09.3	-77:32:08	2.80		0.62	-0.02

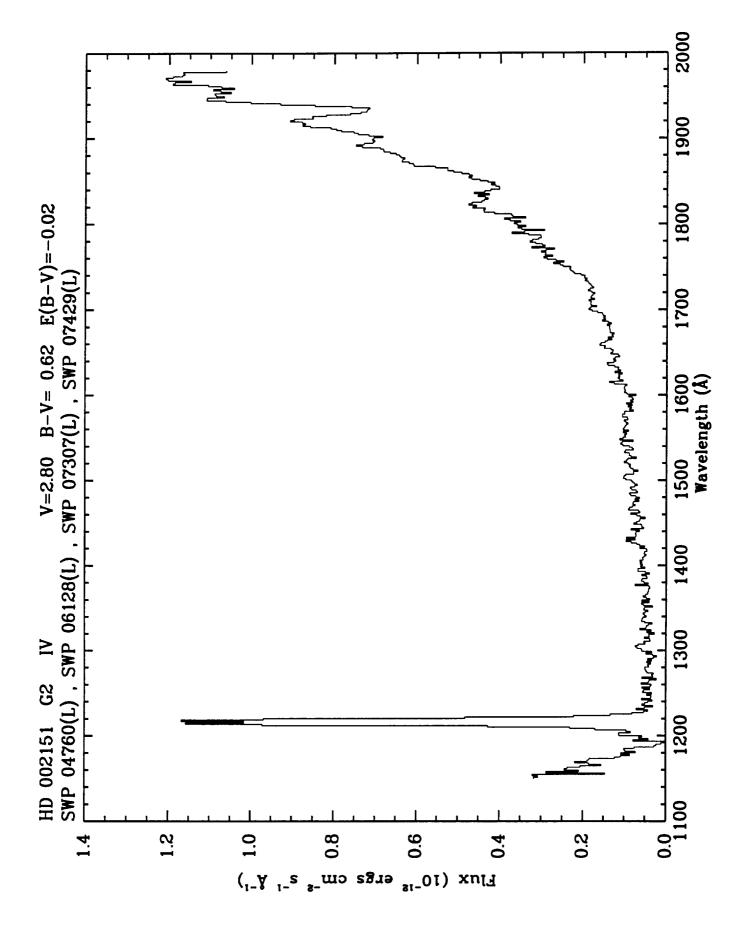
* Low-Dispersion Merged Extracted Image (MXLO)

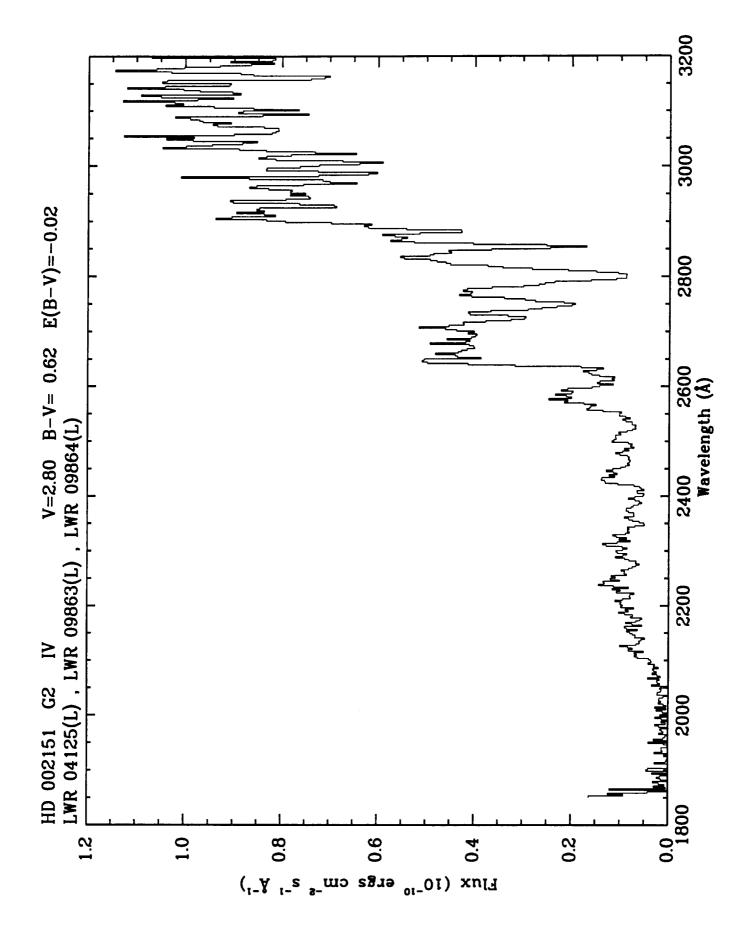
Image	Ap	N	Exp	E/C/B
SWP04760(gif)(ps)(w_f)	L	т	673.535	MAXDN=105
	s	1	599.524	MAXDN=125
SWP06128(gif)(ps)(w_f)	L	1	1019.774	/1-5X/55
	S	1	179.684	/90/60
SWP07307(gif)(ps)(w_f)	L	1	4799.563	/8X/45
	S	1	599.524	/120/45
SWP07429(gif)(ps)(w_f)	L	1	719.537	229/180/25
LWR04125(gif)(ps)(w_f)	L	Т	22.567	MAXDN=265
	S	1	14.620	MAXDN=1.5-2X
LWR09863(gif)(ps)(w_f)	L	Т	18.049	/255/25
LWR09864(gif)(ps)(w_f)	L	Т	16.073	/245/25

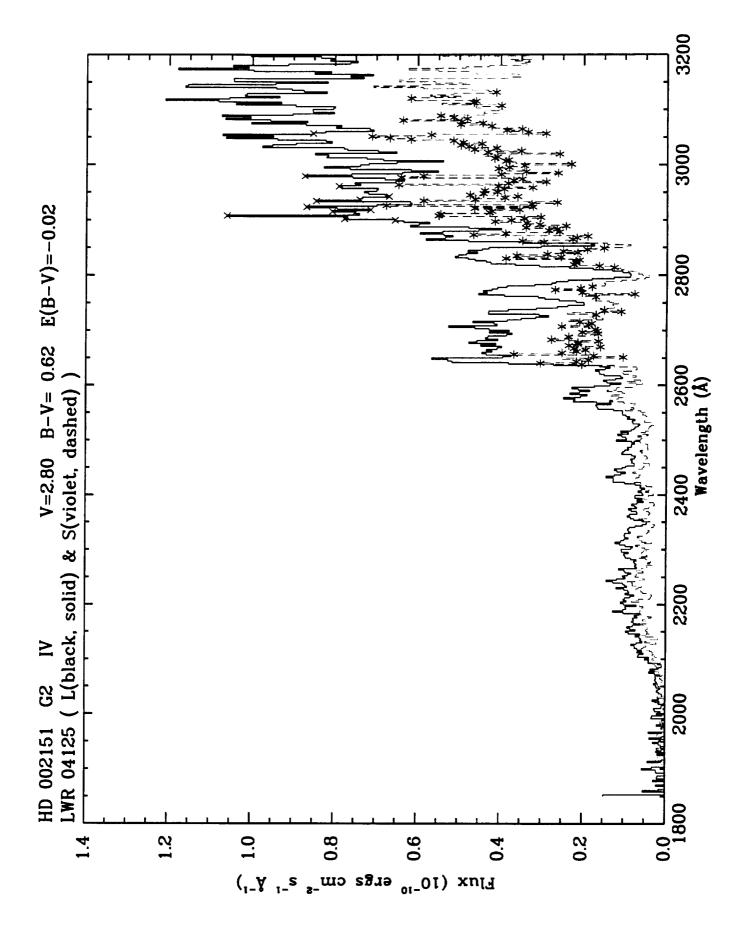
* Combined Image

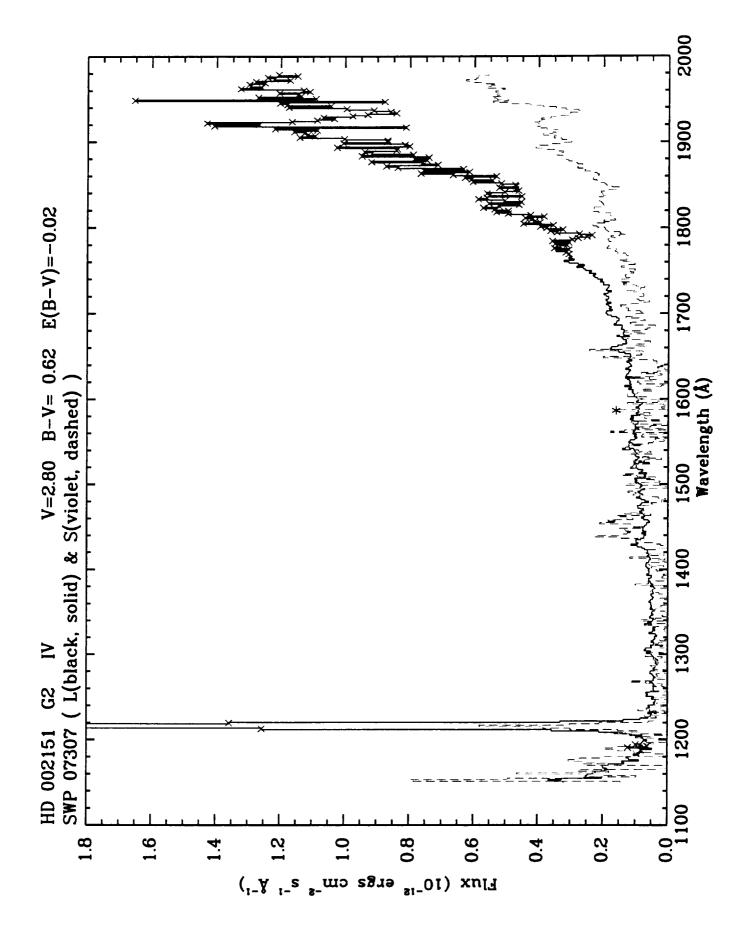
SWP04760, SWP06128, SWP07307, SWP07429(gif)(ps)(w_f) LWR04125, LWR09863, LWR09864 (gif)(ps)(w_f)

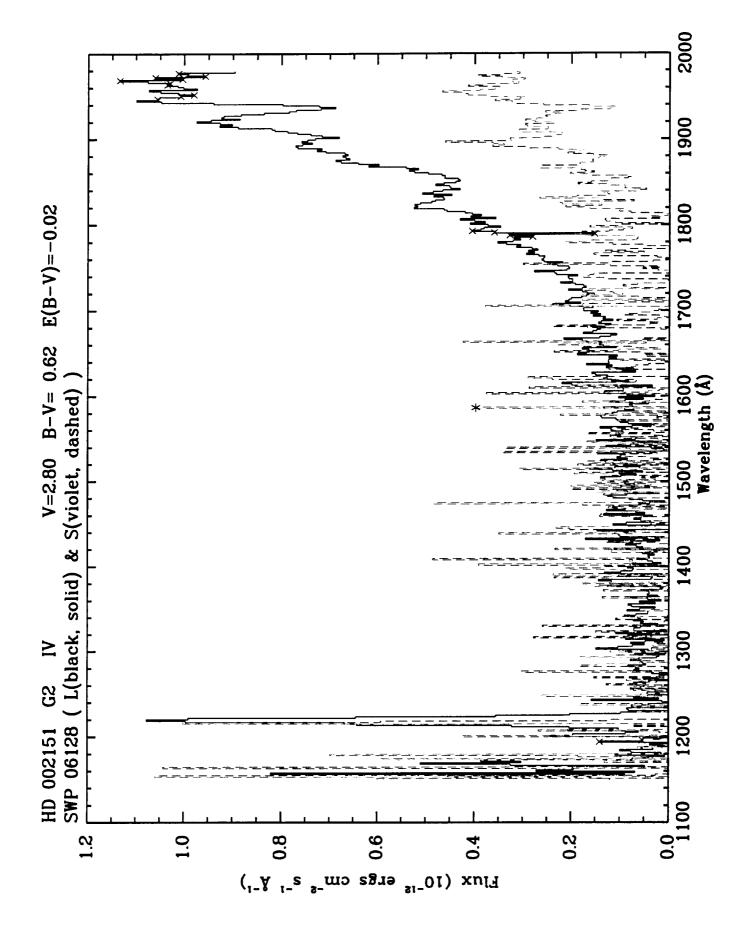
UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	•

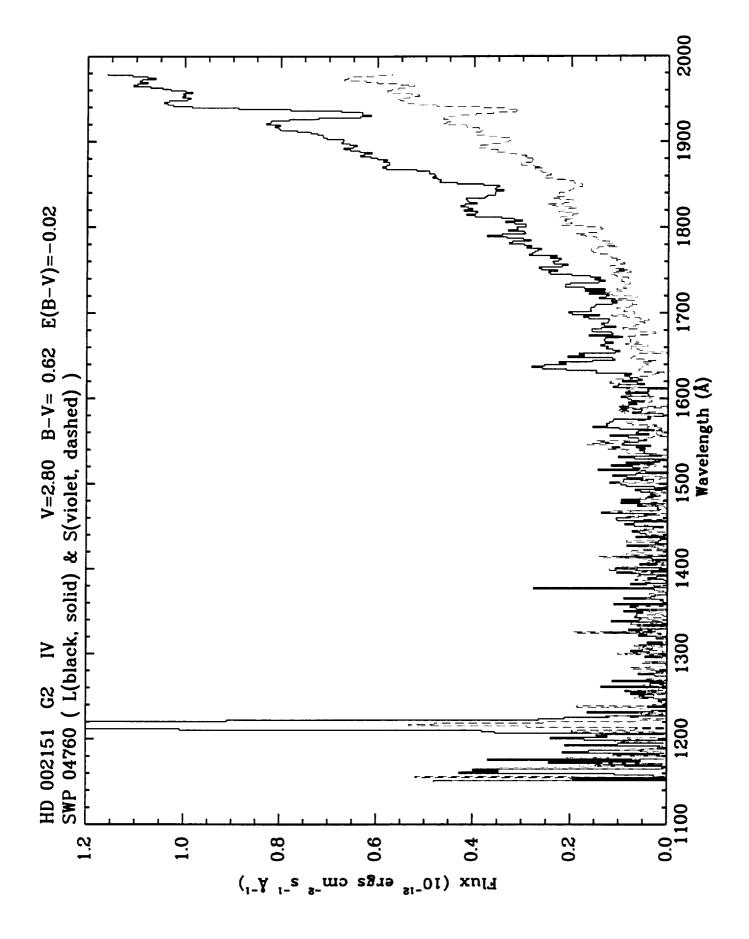












IUE Standard Stars Atlas (NEWSIPS)

HD 62509 BET GEM

Sp_Ty	тре	RA	DEC	v	Rem	B-V	E(B-V)
K0	IIIb	7:42:15.5	28:08:55	1.14	AC	1.00	-0.01

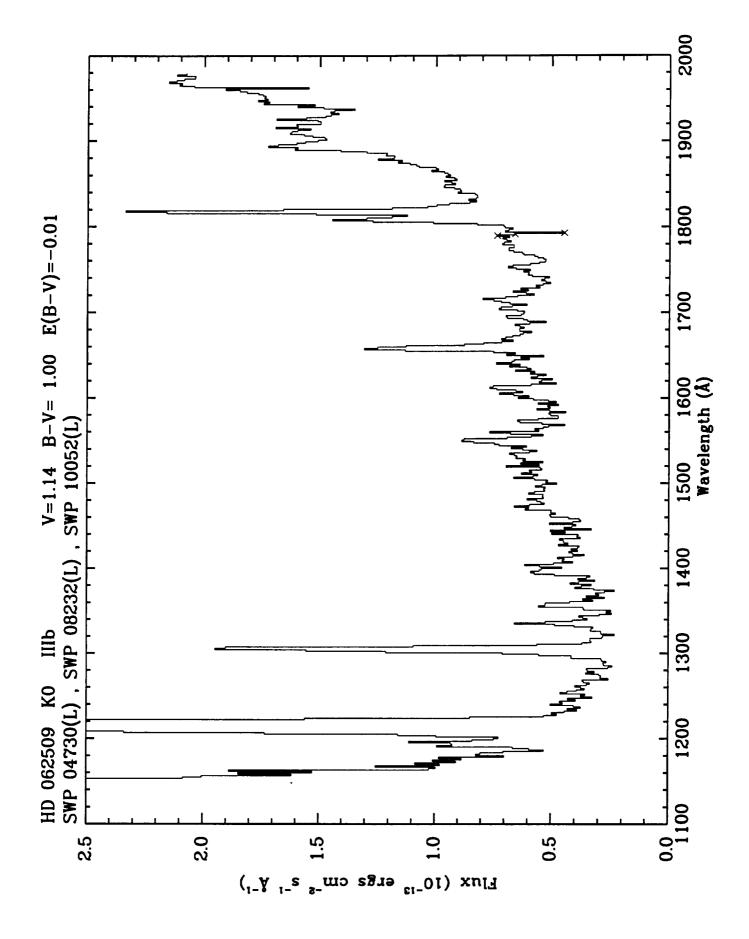
* Low-Dispersion Merged Extracted Image (MXLO)

Image	Ap	N	Ехр	E/C/B
SWP04730(gif)(ps)(w_f)	L	1	1199.588	MAXDN=100
SWP08232(gif)(ps)(w_f)	L	1	7199.819	1.5X/2X/40
	S	1	1499.825	83/110/40
SWP10052(gif)(ps)(w_f)	L	1	2399.717	1.75
LWR09843(gif)(ps)(w_f)	L	\mathbf{T}	6.766	/110/28
LWR09844(gif)(ps)(w_f)	L	\mathbf{T}	15.785	/200/25
LWR09845(gif)(ps)(w_f)	L	\mathbf{T}	132.894	/3X/25

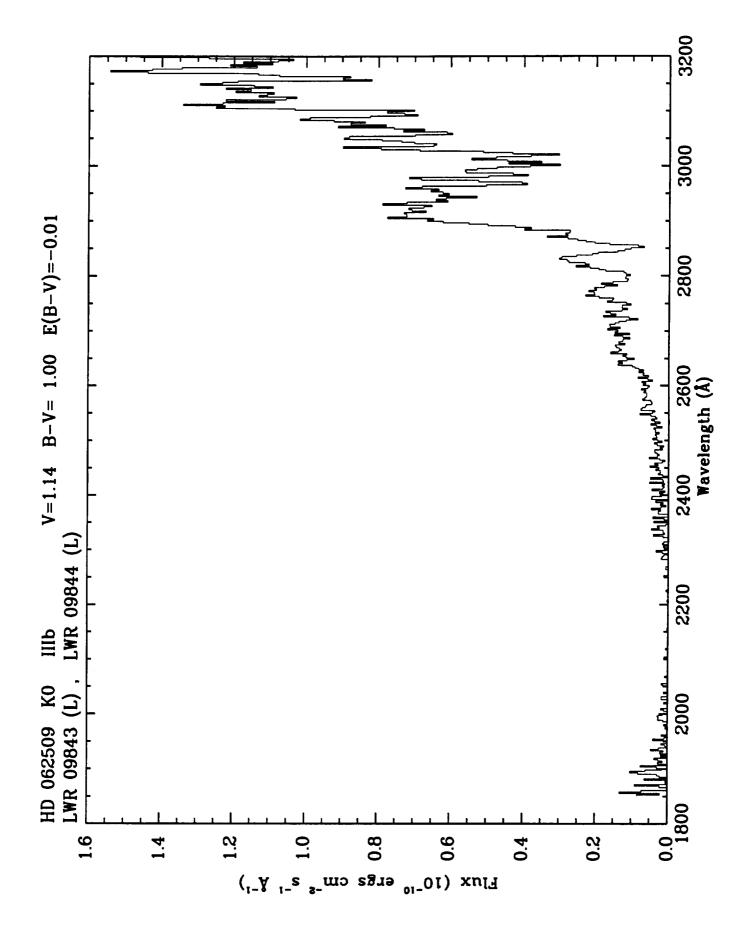
* Combined Image

SWP04730,SWP08232,SWP10052 (gif)(ps)(w_f) LWR09843,LWR09844 (gif)(ps)(w_f)

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UV Atlas Home	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	



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IUE Standard Stars Atlas (NEWSIPS)

HD 146051 DEL OPH

 $\label{eq:sp_Type} \textbf{RA} \qquad \textbf{DEC} \qquad \textbf{V} \qquad \textbf{Rem} \quad \textbf{B-V} \quad \textbf{E}(\textbf{B-V})$

MO.5 III 16:11:43.3 -3:34:01 2.73 A 1.58 -0.01

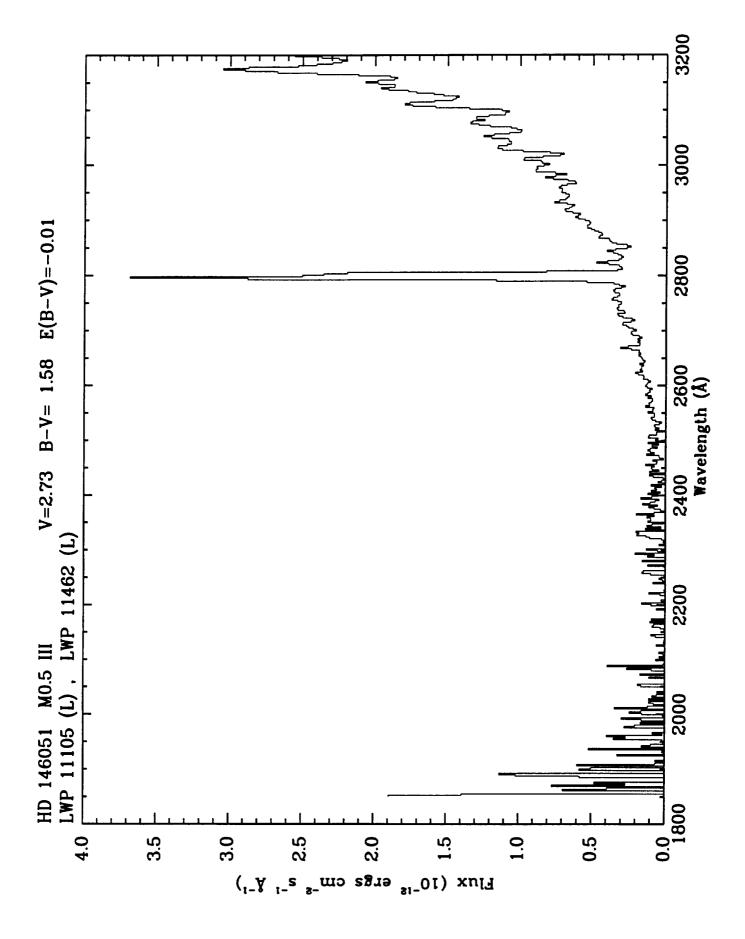
* Low-Dispersion Merged Extracted Image (MXLO)

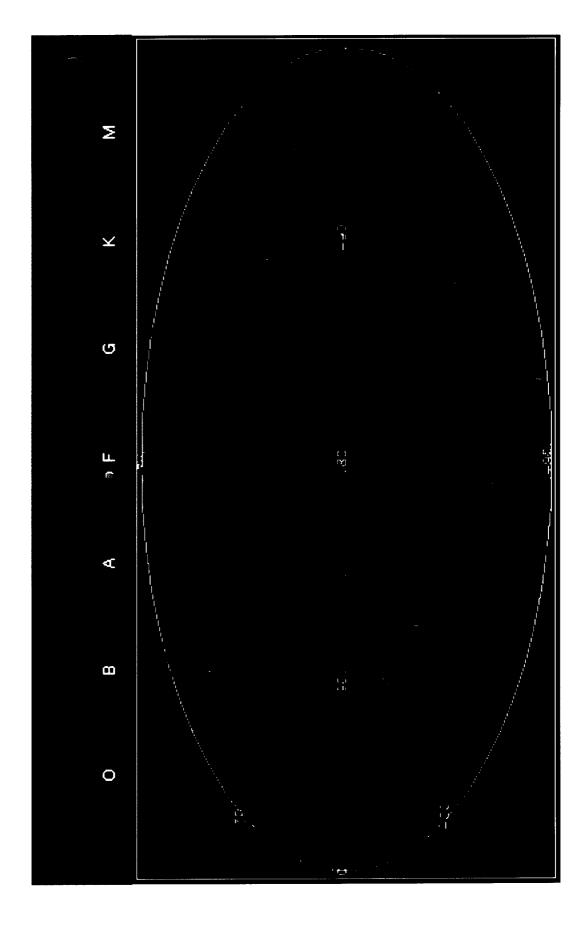
Image		N	Exp	E/C/B	
LWP11105(gif)(ps)(w_f)	_	T	366.507	1.2X/102/34	
LWP11462(gif)(ps)(w_f)		T	220.516	195/75/35	

* Combined Image

LWP11105, LWP11462 (gif) (ps) (w_f)

UV Atlas H	ome l	Project Info	NEWSIPS Atlas	IUESIPS Atlas	IUE Links	•





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